

A Scanning Target Profile Monitor in the Slow Extracted Beam at the AGS

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Introduction

Motivation:

Ever increasing proton beam intensity demands.
Booster injects at higher energy (2GeV vs. 200MeV)
Now using full AGS acceptance.

As intensity exceeded 1×10^{13} switchyard design
Beam size increased.
Causing more halo & losses
At extraction, beam 2X bigger than before the Booster
Resonant, slow extraction process.
150m transport to fixed targets

Existing diagnostics: BLM's suffered from high background
Phosphor screens/video lacked dynamic range

Device: Prototype Scanning target

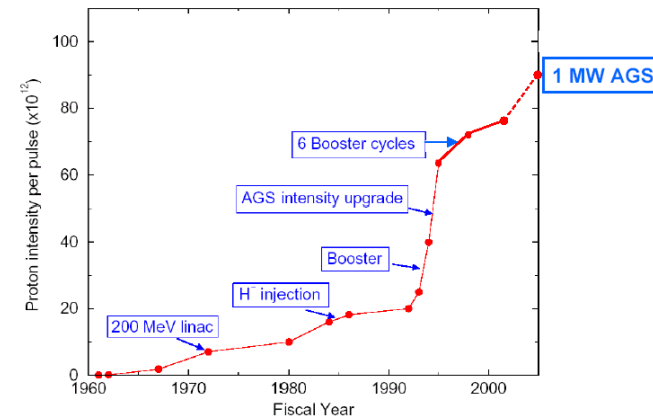
- Scattering into two scintillator telescopes, triple coincidence
- Secondary emission from thin targets
- Targets are movable over the entire 4" aperture

Purpose:

Routinely measured slow extracted beam emittance.

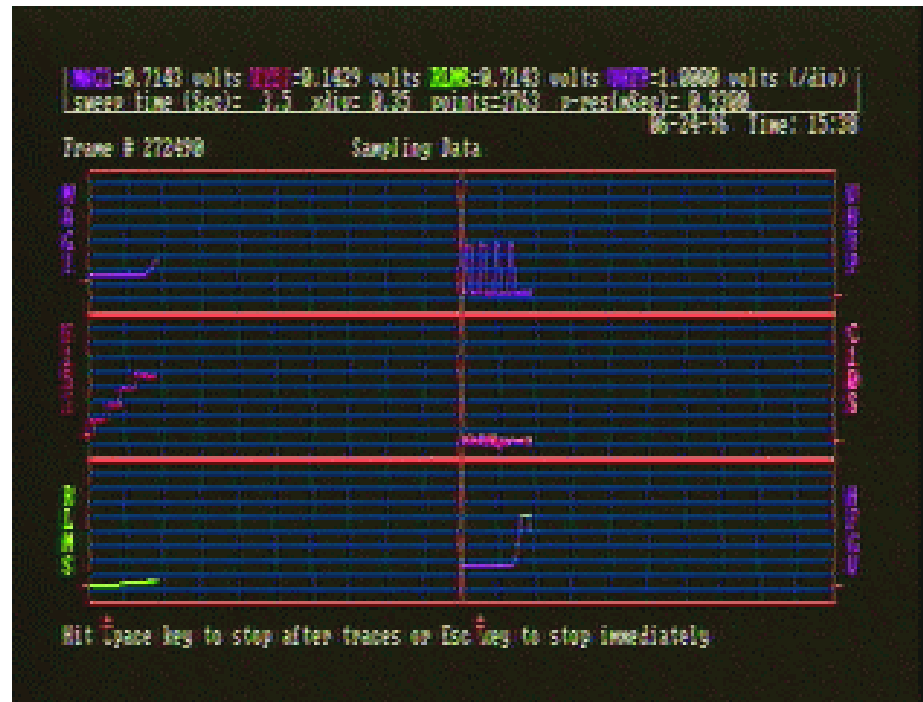
Diagnostic for probing the wings of the beam distribution.

AGS Intensity History



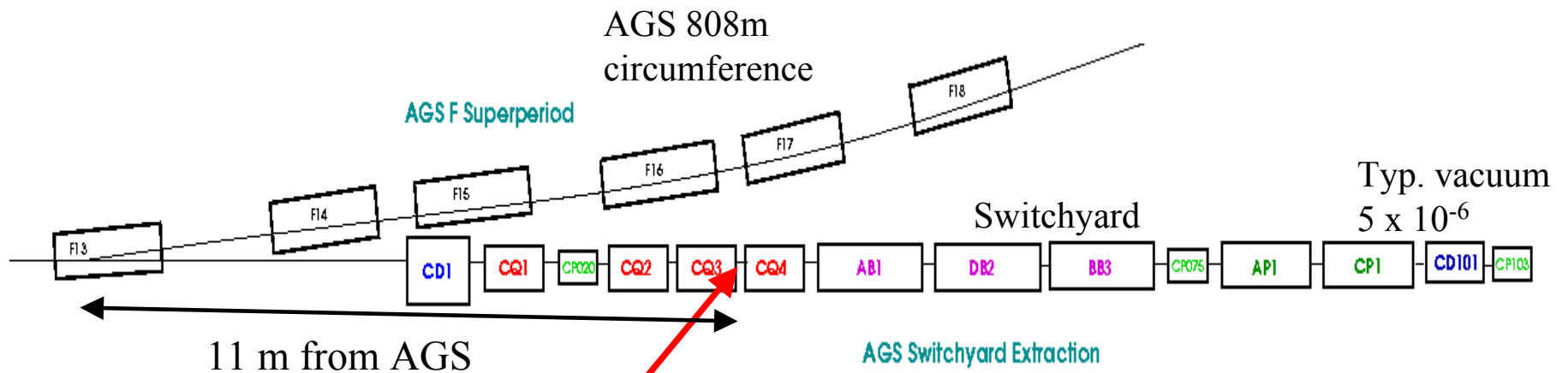
World record proton synchrotron intensity!

Machine Cycle



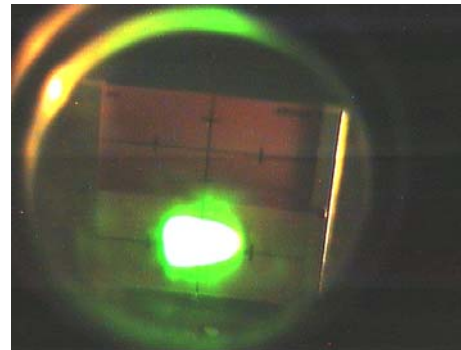
60×10^{12} , one second spill = 10uA average beam-on current

Slow Extracted Beam Transport Layout

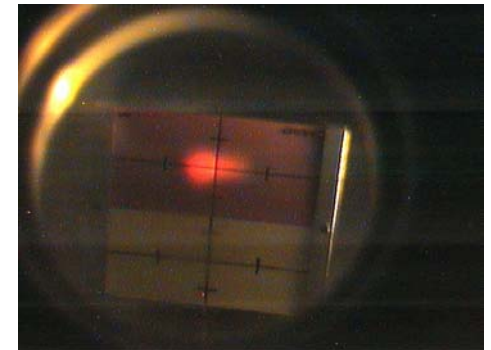


Scanning Target
installed here

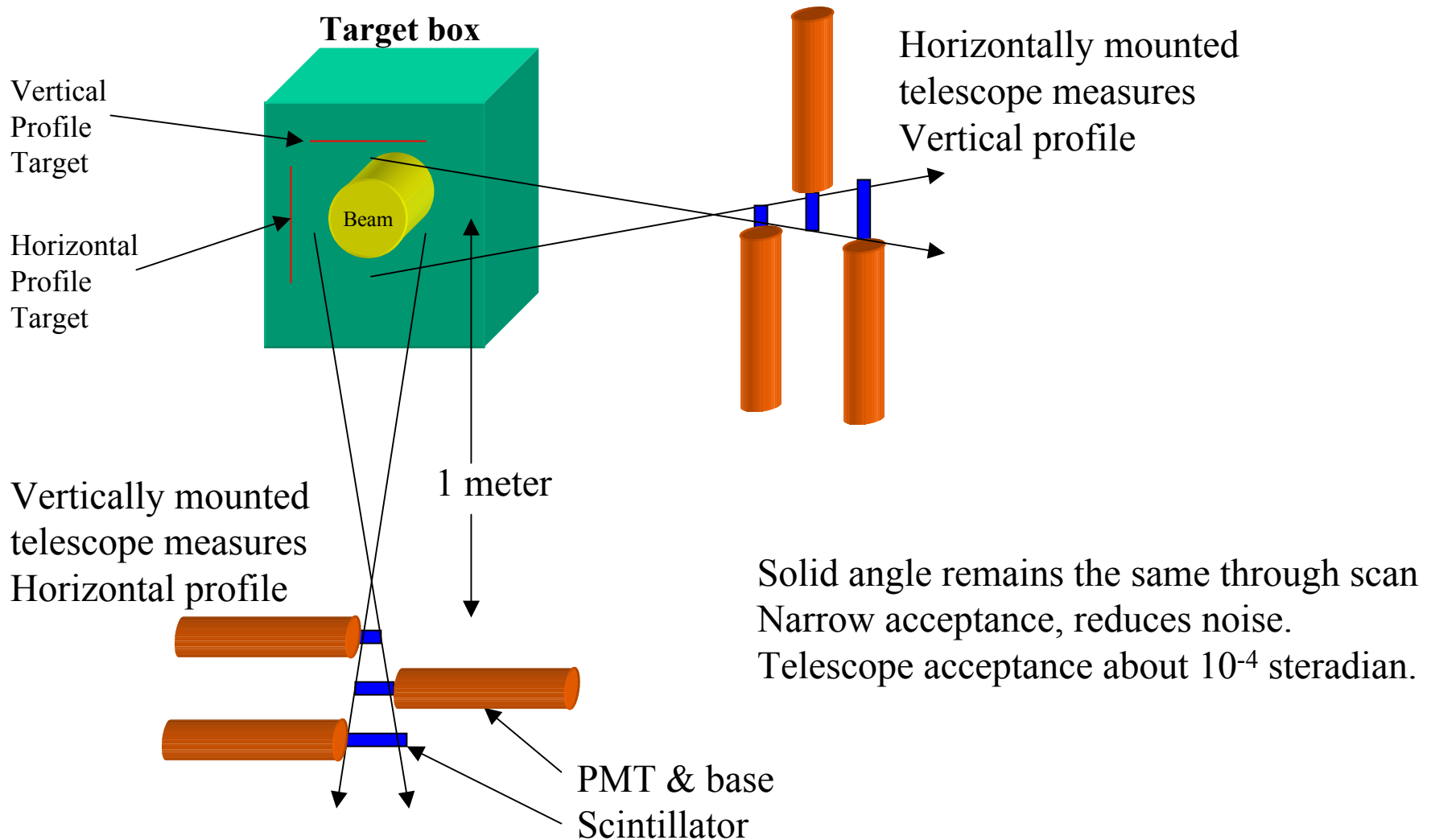
Radlin screen



AlOx screen



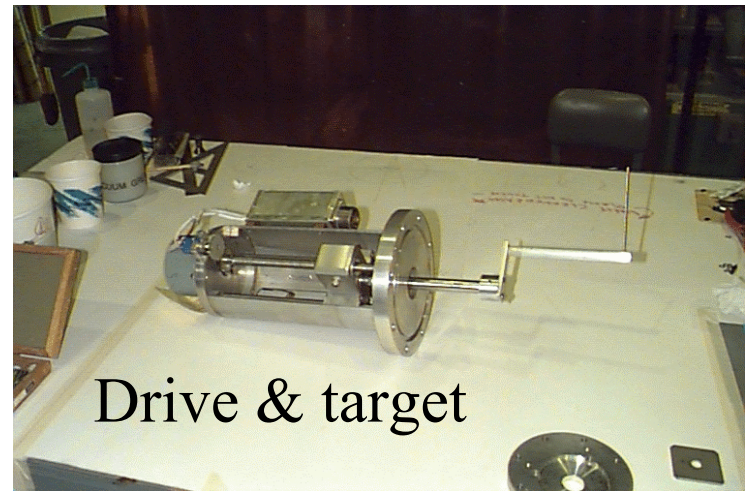
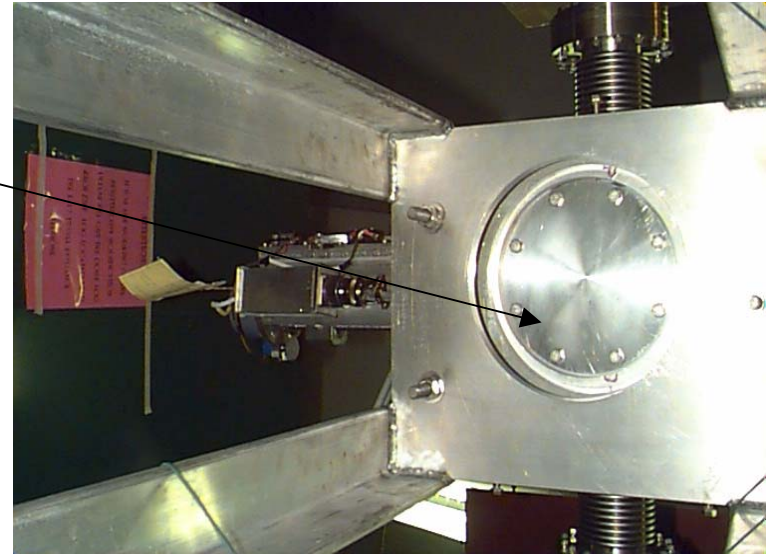
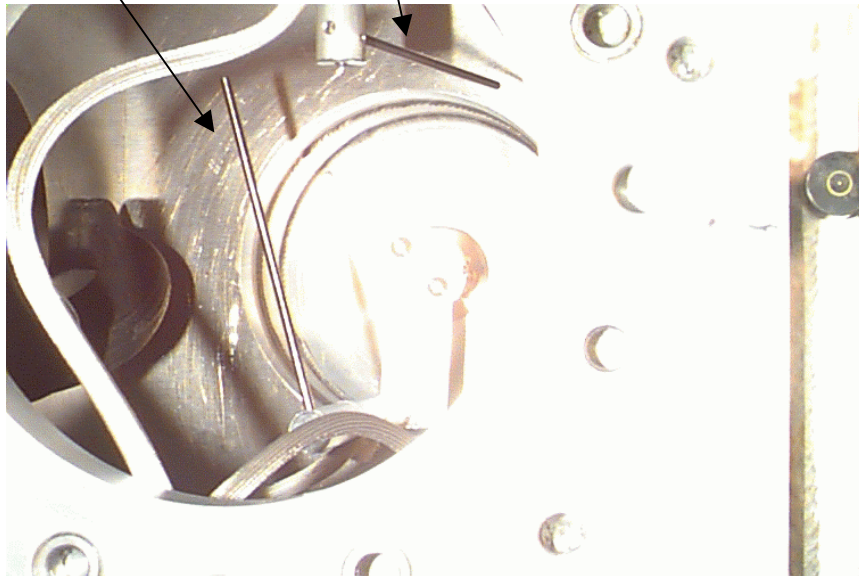
Telescope Operation



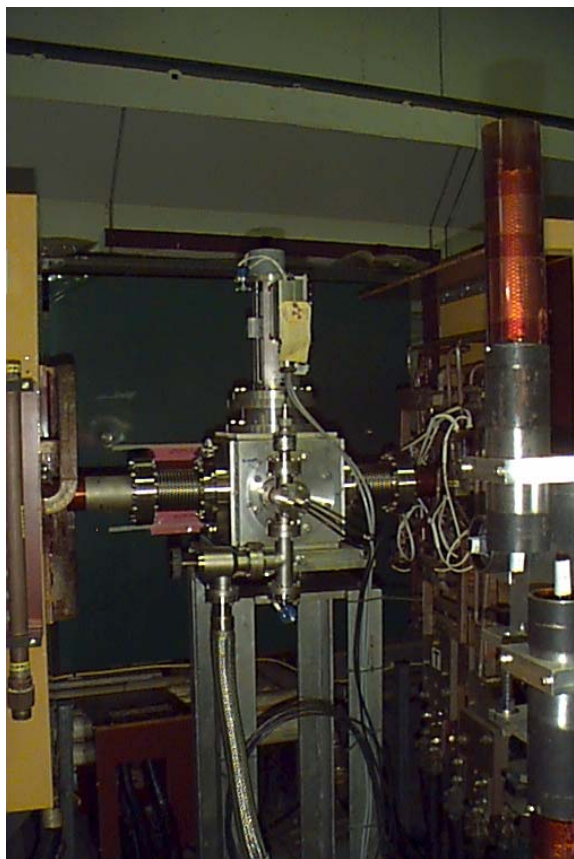
Target Box

Thin Aluminum ports for
 90° scattering.

Tungsten targets, 2.5mm

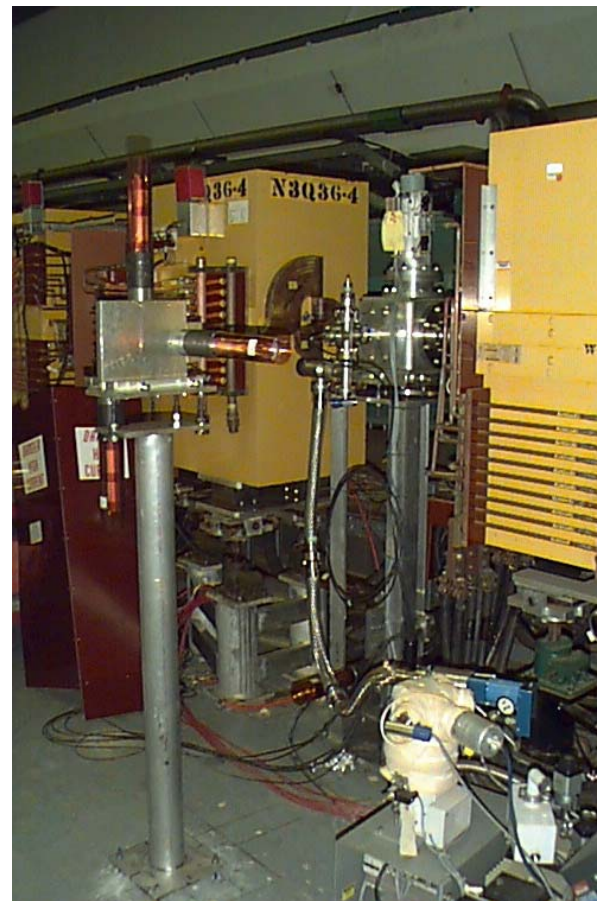


Beamline Assembly



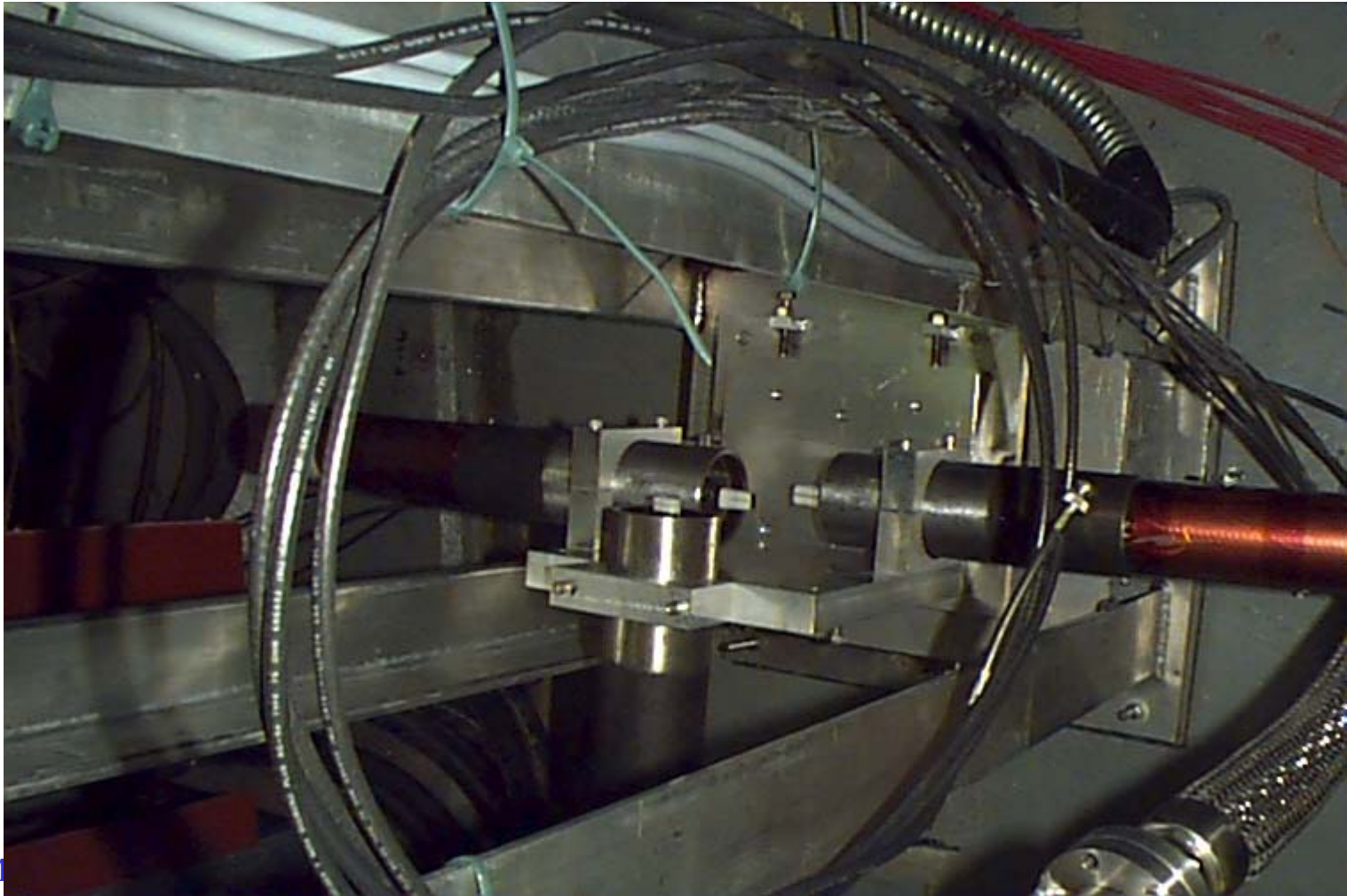
May 21, 2003

Halo '03

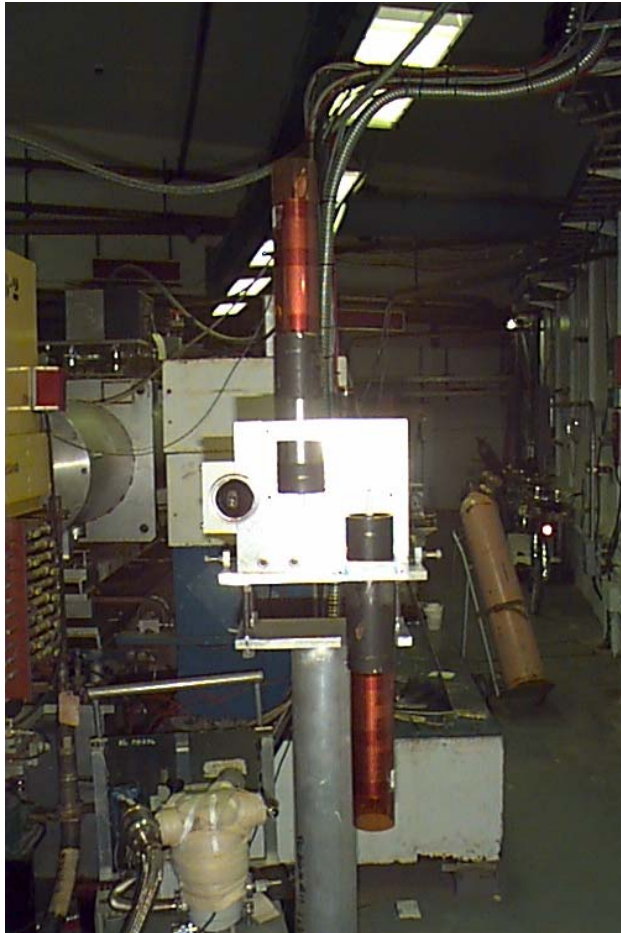


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Horizontal Telescope

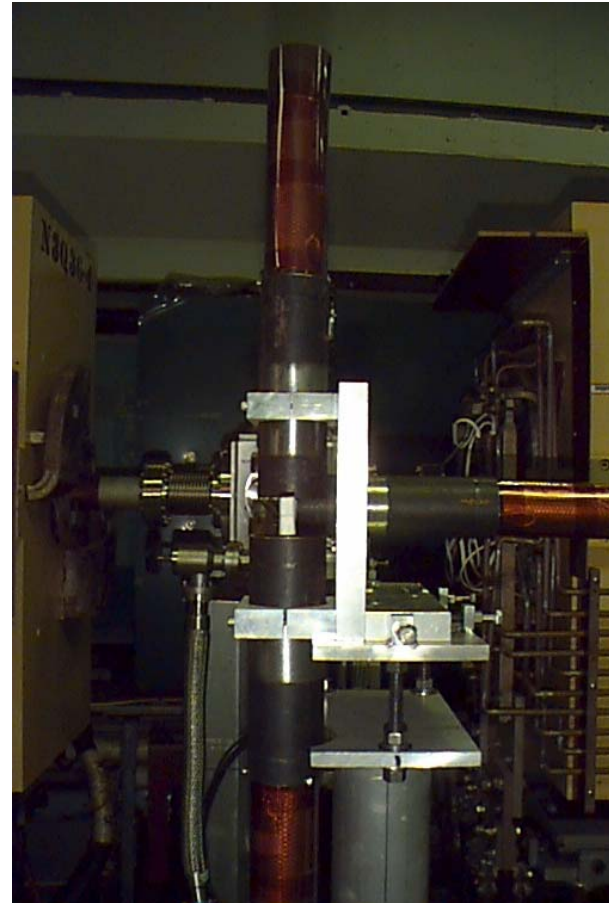


Vertical Telescope



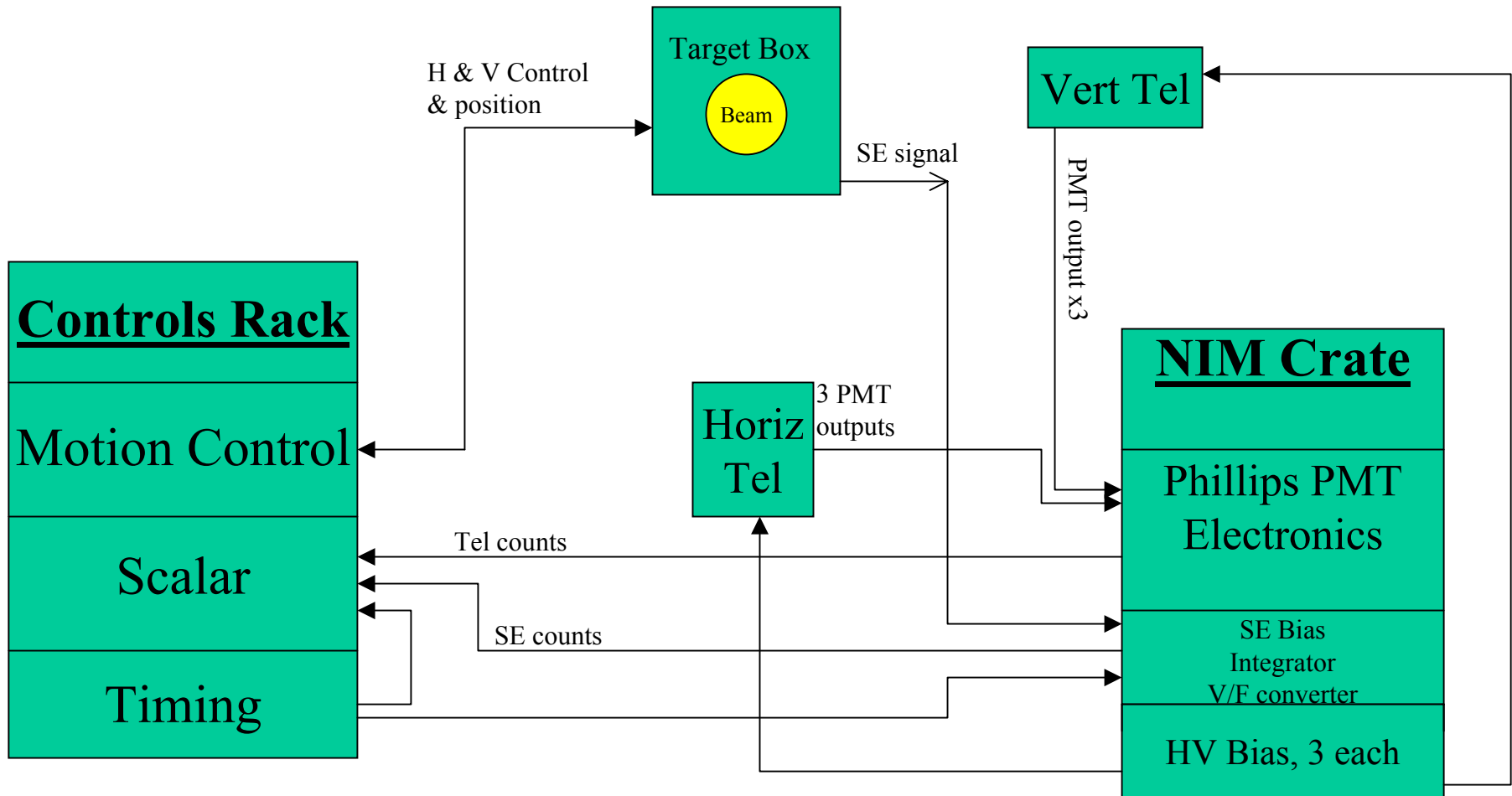
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Halo '03



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System Block Diagram



System Components

Telescope

- Scintillator - Bicron BC404
- PMT - Electron Tubes, Inc. 9813Kb
- BNL designed base, 14 stage, passive
 - Typical pulse FWHM 10ns
- Mu-metal shield, stray fields <10 gauss

NIM Electronics

- Phillips Scientific
 - 770 Amplifier x10
 - 704 Discriminator
 - 754 Coincidence
 - 726 Level converter
- Scalar

Secondary Emission

Isolated, biased target

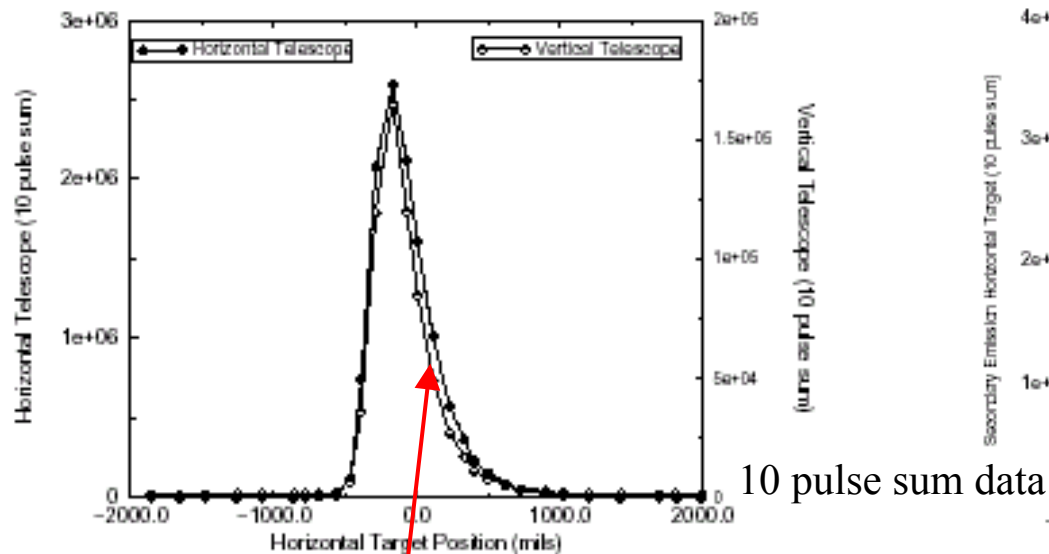
Gated integrator, 1000pF

V/F Converter, 1000 Counts/V

Scalar

Telescope & Secondary Emission Data

Telescope Triples



Secondary Emission data

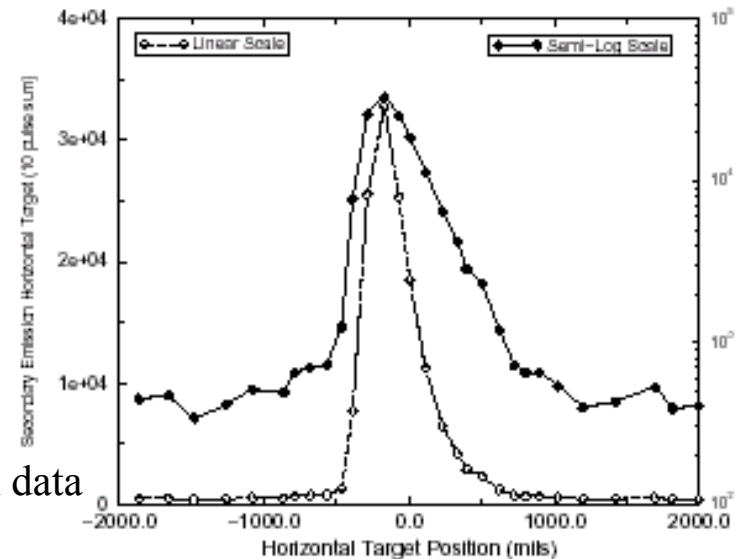


Figure 1: Horiz. and Vert. Telescope triples for Horz. scan

Figure 3: Secondary Emission from target for Horz. Scan

Error by viewing wrong telescope

Estimate beam on target intensity at profile peak, SEM signal.

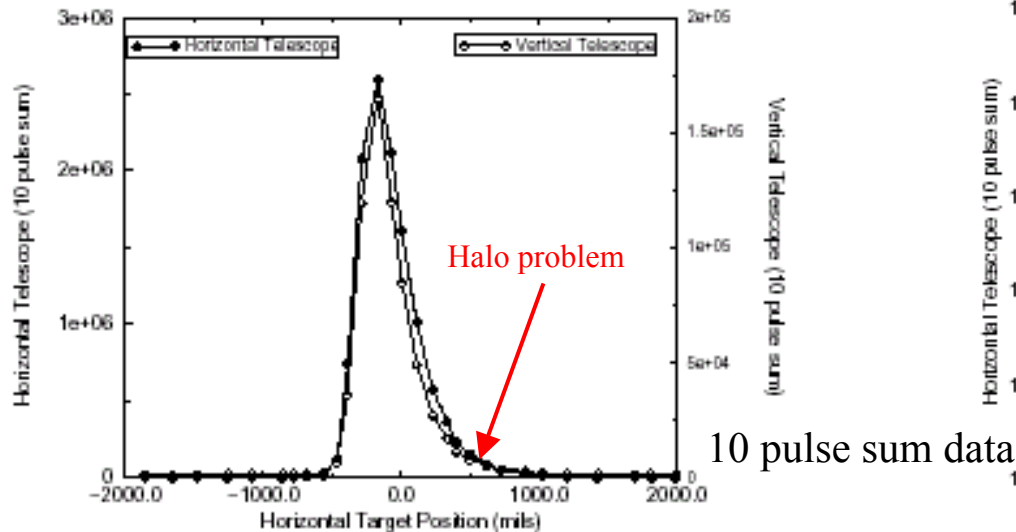
3V on 1nF = 3nA (1 sec spill) X 25(SE%) = 75nA

5e11 protons/sec = 80nA

Beam current 2.5e12 = 400nA

Horizontal Scan, Telescope Data

Linear scale



FWHM approx 1.2cm

Figure 1: Horiz. and Vert. Telescope triples for Horz. scan

Log scale

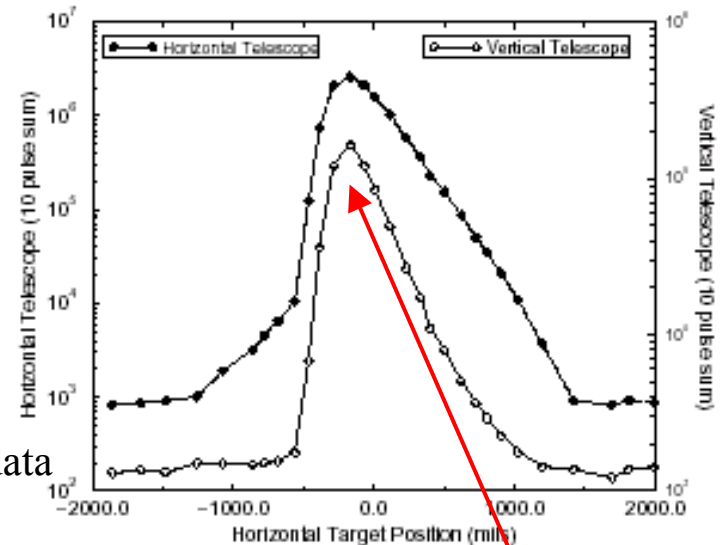


Figure 2: Horiz. and Vert. Telescope triples for Horz. Scan

Measurement shown above, about 10 times below saturation threshold.

S/N 66dB, can resolve 0.05%, DR > 10^3 to 10^4

| | |
|-------------------------|-------|
| Beam current 2.5e12/sec | 400nA |
| Noise floor, 1k | |
| Peak 200k, 5e11 | 75nA |
| +500mils 10k, 2.5e10 | 4nA |
| +1000mils 2k, 2.5e9 | 0.2nA |

Singles rates at peak 5M/sec

Results

Table 1: Summary of emittance measurement results.

(note: β and α are referred to start of SEB line)

| | $\epsilon_x^{95\%,N}$ | β_x (m) | α_x | $\epsilon_y^{95\%,N}$ | β_y (m) | α_y |
|------|-----------------------|------------------|----------------|-----------------------|------------------|----------------|
| FY82 | 31.9 | 57.6 | -6.6 | 38.8 | 3.25 | 0.87 |
| FY96 | 64.4 ± 9.60 | 8.8 ± 1.4 | -0.9 ± 0.2 | 54.7 ± 5.0 | 4.2 ± 0.4 | 1.0 ± 0.09 |

Measurements show a factor of 2 increase as a result of Booster injection.

Halo reduction and understanding gets more critical as intensity increases.

Summary

Scanning target profile monitor proved useful.

Halo problem solved by positioning the extraction kicker and ejection septum further to the inside of the AGS, causing the beam to spend less time in the fringe and edge fields of the AGS main magnets.

Scanning target has since been replaced with current transformer.

Beam profile measurement dynamic range:

| | |
|------------|---------------|
| Telescopes | 4 to 5 orders |
|------------|---------------|

| | |
|-----|---------------|
| SEM | 2 to 3 orders |
|-----|---------------|

Good agreement

Intensity limitation on measurement.

At higher intensities see evidence of saturation.

- high singles rates $> 40\text{M/sec}$

- profile flattening

Improve performance by optimizing (reducing):

Target size

Solid angle acceptance of telescope